# EISENHART-DUVAL LIFT: UNIFIED FRAMEWORK for GALILEI & CARROLL SYMMETRY

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PallaFest Budapest, June 2018 <u>Abstract</u>: While the usual Wigner-Inönü contraction  $c \rightarrow \infty$  of the Poincaré group yields the Galilei group, another  $c \rightarrow 0$  contraction yields the "Carroll group" of Lévy-Leblond. Both boost-invariant theories are conveniently unified within the "Eisenhart-Duval" framework. **Plane gravitational waves** carry a non-trivially implemented "distorted" Carroll symmetry with broken rotations.

Based on:

- C. Duval, G. W. Gibbons, PH and P. M. Zhang *"Carroll versus Newton and Galilei: two dual non- Einsteinian concepts of time,"* Class. Quant. Grav. **31** (2014) 085016 [arXiv:1402.0657 [gr-qc]]
- C. Duval, G. W. Gibbons, PH and P. M. Zhang "Carroll symmetry of gravitational plane waves," Class. Quant. Grav. 34 (2017) 175003 [arXiv:1702.08284]
- P. M. Zhang, C. Duval, G. W. Gibbons and PH, "Velocity Memory Effect for Polarized Gravitational Waves," JCAP 05 (2018) 030 [arXiv:1802.09061 [gr-qc]].
- M. Elbistan, G. W. Gibbons, PH, P.-M. Zhang, "Sturm-Liouville and Carroll : at the heart of the Memory Effect" [arXiv:1803.09640 [gr-qc]]

# **Carroll group**



J. M. Lévy-Leblond

\*

Lewis Carroll

Carroll group \*

constructed as novel type of contraction of Poincaré group

"Une nouvelle limite non-relativiste du group de Poincaré," Ann. Inst. H. Poincaré **3** (1965) 1

V. D. Sen Gupta, "On an Analogue of the Galileo Group," Il Nuovo Cimento **44** (1966) 512

no motion - no physics - mathematical curiosity



"The Red Queen has to run faster and faster in order to keep still where she is. That is exactly what you all are doing!"

Through the

Looking Glass and what Alice Found There (1871).

### NEWTON-CARTAN STRUCTURE

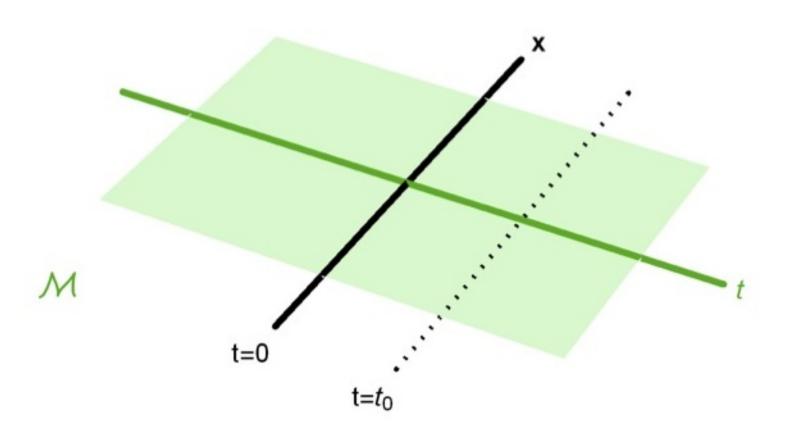


Fig. 1 : Galilean space-time,  $\mathcal{M}$ , described by  $\begin{pmatrix} x \\ t \end{pmatrix}$ . Carries symmetric, contravariant non-negative [spaceco-] "metric" tensor  $\gamma$  whose kernel is generated by dt. Projects onto absolute time.

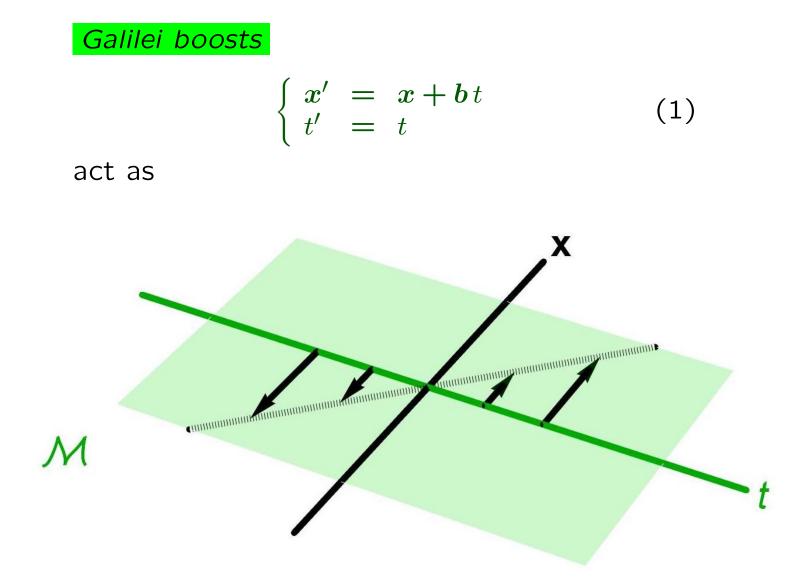


Fig.2 Galilei boost acting on Galilei space-time

#### CARROLL STRUCTURE

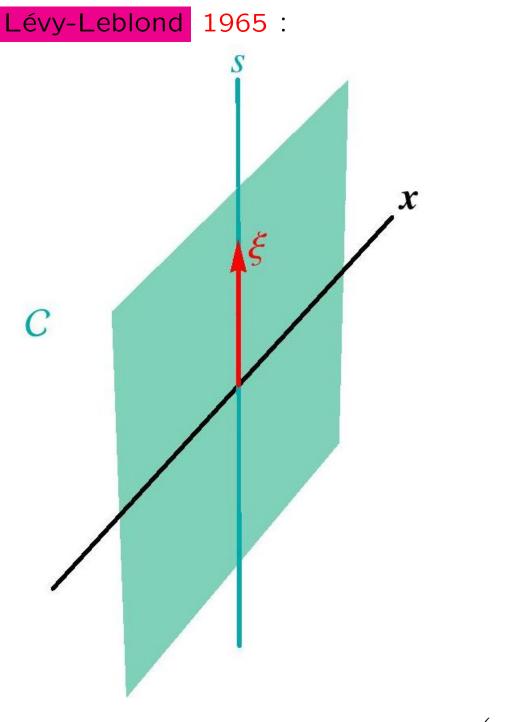


Fig.2 : Carroll space-time, C described by  $\begin{pmatrix} x \\ s \end{pmatrix}$ , is endowed with vector  $\boldsymbol{\xi}$  which generates kernel of (singular) [space-] "metric"  $\bar{G} = \delta_{AB} dx^A dx^B$ .

Carroll group  $Carr(d+1) \sim$ 

"Carrollian boosts"

$$\begin{cases} x' = x \\ s' = s - b \cdot x \end{cases}$$
(2)

NB: In NR QM wave fct including a phase factor,

$$\psi'(x,t) = e^{i(\mathbf{b}\cdot\boldsymbol{x} - \frac{1}{2}\mathbf{b}^2t)}\psi(x - \mathbf{b}t, t)$$
(3)

infinitesimally (for t = 0):

$$\widehat{\boldsymbol{\beta}}\psi(\boldsymbol{x},0) = (-\boldsymbol{\beta}\cdot\boldsymbol{x})\psi(\boldsymbol{x},0)$$
 (4)

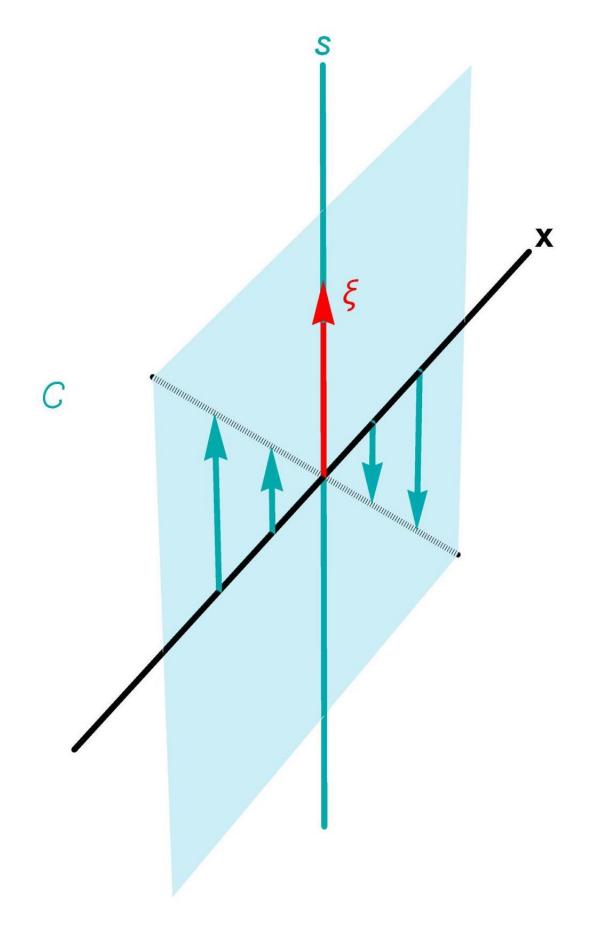


Fig.3 Carroll boost acts on flat Carroll space-time

Carroll group represented by  $(d+2) \times (d+2)$  matrices

$$\begin{array}{ccc} R & 0 & \mathbf{c} \\ -\mathbf{b}^T R & 1 & f \\ 0 & 0 & 1 \end{array}$$
 (5)

where  $R \in O(d)$ ,  $\mathbf{b}, \mathbf{c} \in \mathbb{R}^d$ ,  $f \in \mathbb{R}$ . Acts on  $\begin{pmatrix} x \\ s \\ 1 \end{pmatrix}$  affinely by matrix action. Carroll Lie algebra  $\operatorname{carr}(d+1)$ 

$$Z = \begin{pmatrix} \omega & 0 & \gamma \\ -\beta^T & 0 & \varphi \\ 0 & 0 & 0 \end{pmatrix}$$
(6)

 $\omega \in \mathfrak{so}(d)$ ,  $\beta, \gamma \in \mathbb{R}^d$ , and  $\varphi \in \mathbb{R}$  acts on Carroll space-time as

$$X = (\omega_B^A x^B + \gamma^A) \frac{\partial}{\partial x^A} + \left(\varphi \left[-\beta_A x^A\right]\right) \frac{\partial}{\partial s}, \quad (7)$$
  
where  $\omega \in \mathfrak{so}(d), \quad \beta, \gamma \in \mathbb{R}^d$ , and  $\varphi \in \mathbb{R}$ .

N.B.: Galilei Lie algebra 
$$\mathfrak{gal} \equiv \mathfrak{gal}(d+1)$$
  
$$X = (\omega_B^A x^B + \beta^A t + \gamma^A) \frac{\partial}{\partial x^A} + \epsilon \frac{\partial}{\partial t} \qquad (8)$$
where  $\omega \in \mathfrak{so}(d), \ \beta, \gamma \in \mathbb{R}^d$  and  $\epsilon \in \mathbb{R}$ .

Unification: Bargmann manifolds

#### A Bargmann manifold\* is

(i) a (d+2)-dim manif B

(ii) endowed with metric G of signature (d + 1, 1)

(iii) carries nowhere vanishing, complete, null "vertical" vector  $\xi$ , parallel-transported by Levi-Civita connection,  $\nabla$ .

L. P. Eisenhart, "Dynamical trajectories and geodesics", Annals. Math. **30** 591-606 (1928).

J. B. Kogut and D. E. Soper, "Quantum Electrodynamics in the Infinite Momentum Frame," Phys. Rev. D 1(1970) 2901.

C. Duval, G. Burdet, H. P. Kunzle and M. Perrin, "Bargmann Structures and Newton-Cartan Theory," Phys. Rev. D **31** (1985) 1841.

\* Introduced by **Duval** et al as geometrical structure underlying **Bargmann** [ $\equiv$  centrally extended Galilei] group.

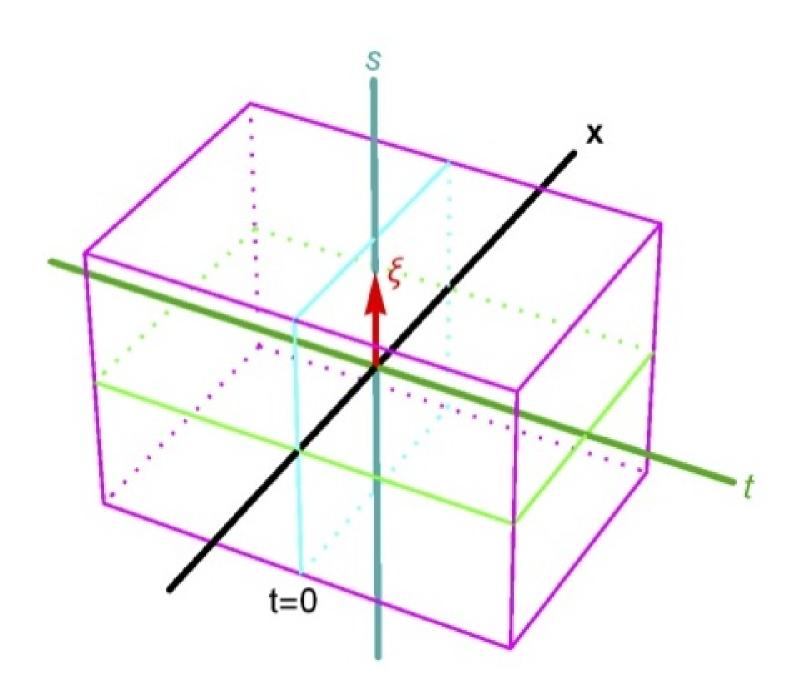


Fig. 4 : Bargmann space : (d+1,1) dim manifold with Lorentz metric & coordinates (x,t,s), endowed with covariantly constant null vector  $\boldsymbol{\xi} = \partial_s$ .

Flat Bargmann structure  $\sim$  Minkowski space :

$$B = \mathbb{R}^d \times \mathbb{R} \times \mathbb{R} = \left\{ \begin{pmatrix} x \\ t \\ s \end{pmatrix} \right\}, \quad (9)$$

$$G = \delta_{AB} dx^A dx^B + 2dt ds, \qquad (10)$$

$$\boldsymbol{\xi} = \partial_s. \tag{11}$$

Both s & t light-cone (null), coords. t has dimension of time, coordinate s has that of action/mass.

• Factoring out "vertical" translations along  $\xi$ , (d+1)-dim quotient acquires Newton-Cartan structure

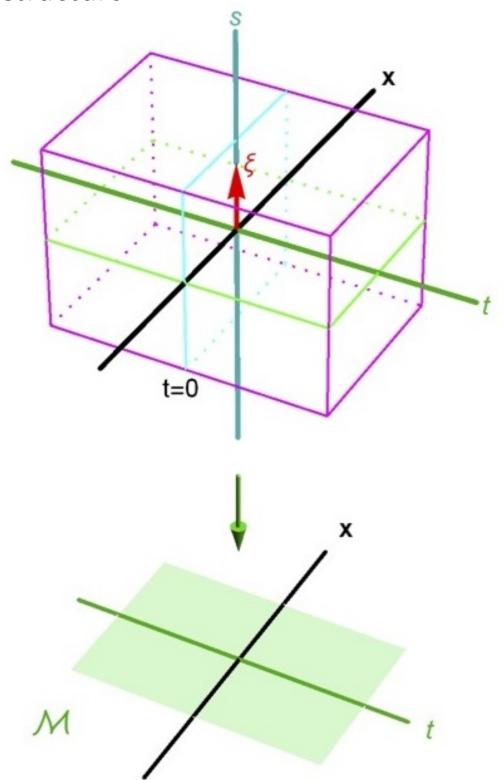


Fig.5 : Bargmann projects to Galilean space-time.

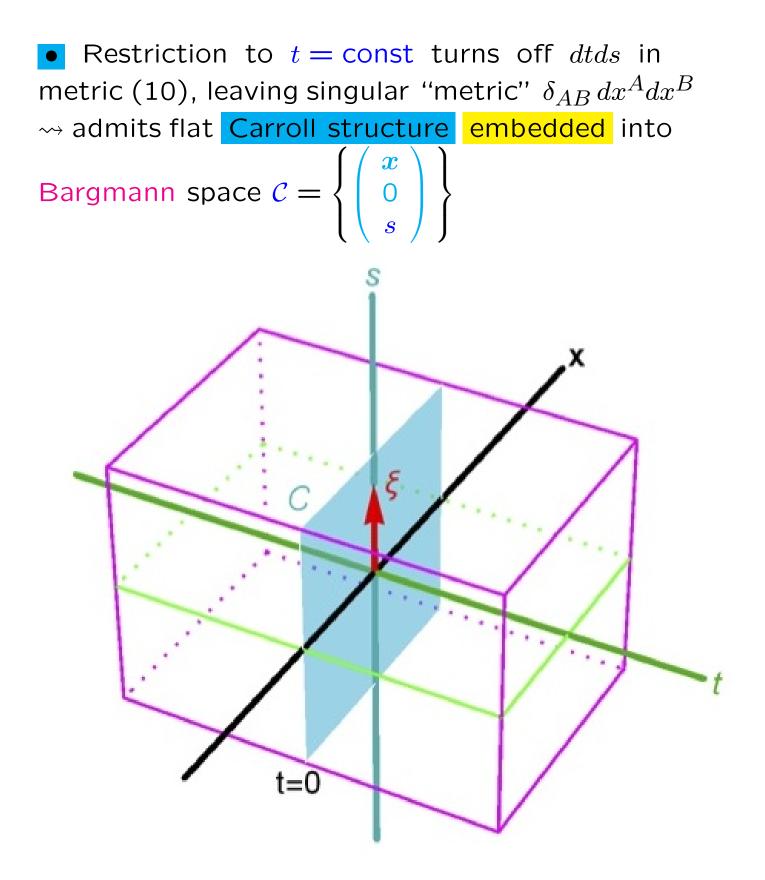


Fig.6 : t = const slice is "Carroll space-time" C embedded into Bargmann space.

# **Symmetries**

 $\xi$ -preserving isometries of Bargmann :

$$a = \begin{pmatrix} R & b & 0 & c \\ 0 & 1 & 0 & e \\ -b^T R & -\frac{1}{2}b^2 & 1 & f \\ 0 & 0 & 0 & 1 \end{pmatrix}$$
(12)

where  $R \in O(d)$ ,  $\mathbf{b}, \mathbf{c} \in \mathbb{R}^d$ , and  $e, f \in \mathbb{R}$  form centrally extended Galilei [ $\equiv$  Bargmann] group **Barg**  $\equiv$  Barg(d+1). Boost :

$$\begin{pmatrix} \boldsymbol{x} \\ t \\ s \end{pmatrix} \rightarrow \begin{pmatrix} \boldsymbol{x} + \mathbf{b}t \\ t \\ s - \mathbf{b} \cdot \boldsymbol{x} + \frac{1}{2}\mathbf{b}^{2}t \end{pmatrix}$$
(13)

**N.B.** : lifting ordinary wave fct to equivariant  $(\equiv \partial_s \Psi = im\Psi)$  on B-space, Galilei boost action (3) is Bargmann action. Affine action on  $\begin{pmatrix} x \\ t \\ s \\ 1 \end{pmatrix} \rightsquigarrow \text{Bargmann algebra } \mathfrak{barg} \equiv \mathfrak{barg}(d+1)$ 

$$(\omega_B^A x^B + \beta^A t + \gamma^A) \frac{\partial}{\partial x^A} + \varepsilon \frac{\partial}{\partial t} + (\varphi - \beta_A x^A) \frac{\partial}{\partial s}$$
(14)

where  $\omega \in \mathfrak{so}(d)$ ,  $\beta, \gamma \in \mathbb{R}^d$ ,  $\varepsilon, \varphi \in \mathbb{R}$ .

Seen before: restriction of Bargmann space to t = 0 is Carroll manifold C left invariant by restriction of Bargmann action (14) with  $e = 0 \rightsquigarrow$  action of Carr, embedded into Bargmann group,

$$\begin{pmatrix} R & 0 & c \\ -b^T R & 1 & f \\ 0 & 0 & 1 \end{pmatrix} \hookrightarrow \begin{pmatrix} R & b & 0 & c \\ 0 & 1 & 0 & 0 \\ -b^T R & -\frac{1}{2}b^2 & 1 & f \\ 0 & 0 & 0 & 1 \end{pmatrix}$$
(15)

where  $R \in O(d)$ ,  $\mathbf{b}, \mathbf{c} \in \mathbb{R}^d$ ,  $f \in \mathbb{R}$ .

Carr(d+1) : e = 0 subgroup of Barg(d+1).

Infinitesimally:

$$(\omega_B^A x^B + \gamma^A) \frac{\partial}{\partial x^A} + (\varphi - \beta_A x^A) \frac{\partial}{\partial s}$$
(16)  
$$\omega \in \mathfrak{so}(d), \ \beta, \ \gamma \in \mathbb{R}^d, \ \varphi \in \mathbb{R} \text{ (seen before).}$$

N.B.: for 
$$t = t_0$$
 Carroll boost acts as  
 $v \rightarrow v - \mathbf{b} \cdot \mathbf{x} - \frac{1}{2}\mathbf{b}^2 t_0$  (17)

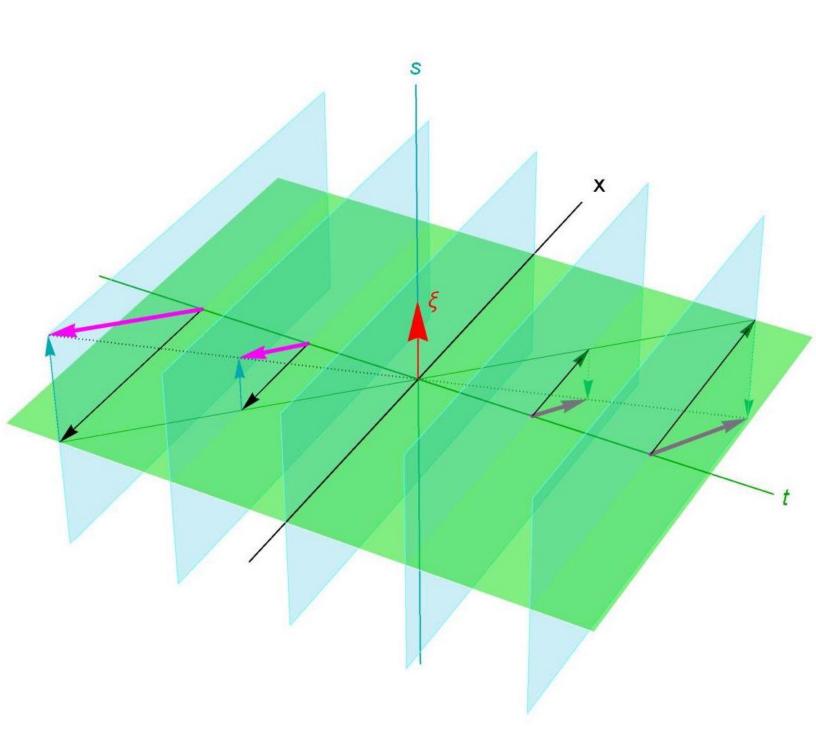


Fig.7 Boost acting on flat Bargmann [ $\equiv$  Minkowski] space

## Plane gravitational waves

#### In Brinkmann coordinates

 $ds^{2} = d\boldsymbol{X}^{2} + 2dUdV - K(U, \boldsymbol{X}) dU^{2} \quad (18)$ 

U and V light-cone coords,  $\boldsymbol{X}=(X_1,X_2)\sim$  transverse plane. Vacuum Einstein eqn satisfied with

 $K(U, \mathbf{X}) = \mathcal{A}(U) \left( X_1^2 - X_2^2 \right) + 2\mathcal{B}(U) X_1 X_2.$ (19)

Clue: (18) Bargmann space  $\sim$  anisotropic oscillator.

P. M. Zhang, P. A. Horvathy, K. Andrzejewski, J. Gonera and P. Kosinski, "Newton-Hooke type symmetry of anisotropic oscillators," Ann. Phys. **333** (2013) 335 [arXiv:1207.2875 [hep-th]].

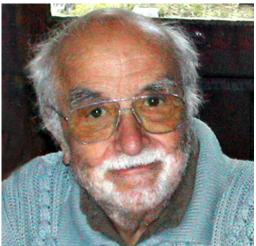
Isometries : Bondi et al 1959. 5-parameters. 1 "vertical" translation + 4 MYSTERIOUS (not written explicitely). Torre 2006 Gen. Rel. Grav. 38 (2006) 653 : isometries  $\partial_V$  +

$$S_i(U)\partial_i + \dot{S}_i(U)X_i\partial_V \tag{20}$$

where  $S_i$  is solution of Sturm-Liouville eqn

$$\ddot{S}_i = K_{ij}(U) S_j \tag{21}$$

(... that we can't solve in general ...)



Souriau 1973 metric in BJR (Baldwin-Jeffery-Rosen) coords :

$$ds^{2} = \boxed{a_{ij}(u)} dx^{i} dx^{j} + 2dudv.$$
 (22)

Isometries :  $u \rightarrow u$ , completed with

 $x \to x + H(u)\mathbf{b} + \mathbf{c},$  (23a)

$$v \rightarrow v - \mathbf{b} \cdot \boldsymbol{x} - \frac{1}{2}\mathbf{b} \cdot H(u)\mathbf{b} + f$$
 (23b)

where  $H = (H_{ij})$  is 2 × 2 matrix

$$H(u) = \int_{u_0}^{u} a^{-1}(w) dw.$$
 (24)

 $\mathbf{c} \in \mathbb{R}^2 \sim \text{transverse-space transl, } f \sim \text{null translat along } v$  coord.

Group composition law: that of Carroll group with no rotations.  $\mathbf{b} \in \mathbb{R}^2$  generates Carroll boost, implemented as in (23).

Flat case: 
$$a_{ij} = \delta_{ij} \Rightarrow$$
  
 $H(u) = (u - u_0) \operatorname{Id}$  (25)  
choosing  $u_0 = 0$   
 $x \to x + u \operatorname{b},$  (26a)  
 $u \to u,$  (26b)

$$v \to v - \mathbf{b} \cdot \boldsymbol{x} - \frac{1}{2}\mathbf{b}^2 u$$
 (26c)

Galilei boosts lifted to flat Bargmann space.

(See again at the end)

Relation with Brinkmann-coords ?

1. Given B-profile K(U), solve Sturm-Liouville

$$\ddot{P}_{kj} = K_{kr} P_{rj} \tag{27}$$

for U-dept 2 × 2 matrix  $P_{kj}(U)$  .

2. Putting

$$X^{i} = P_{ij} x^{j} \qquad U = u$$
(28a)  
$$a_{ij}(u) = P_{ri} P_{rj}, \quad V = v - \frac{1}{4} \frac{da_{ij}}{du} x^{i} x^{j}$$
(28b)

allows to present metric (18) in BJR form

$$ds^2 = \boxed{a_{ij}(u)} dx^i dx^j + 2dudv$$

cf. (22) provided also  $P^{\dagger}\dot{P} = \dot{P^{\dagger}}P$ .

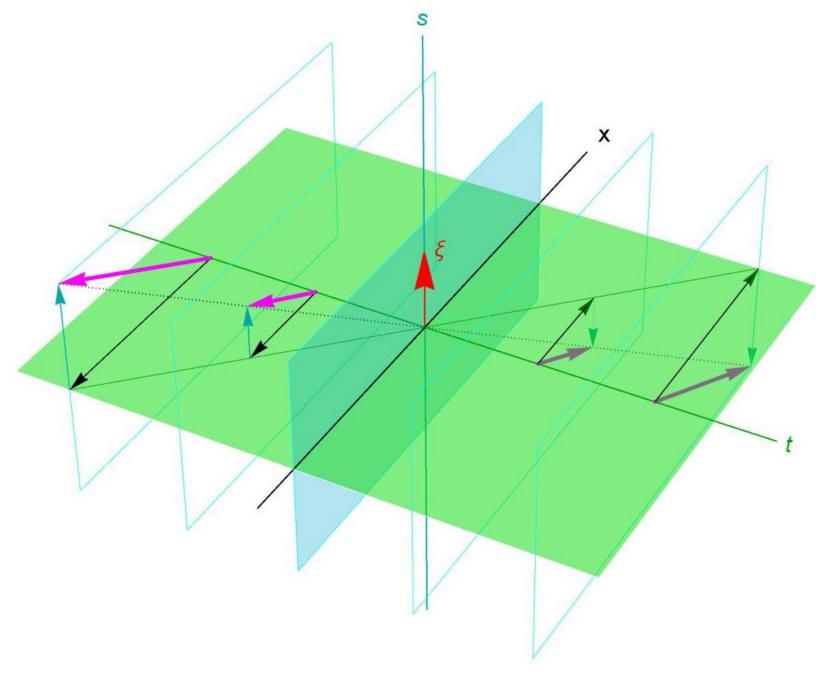
Quadratic "scalar potential" in B,  $K_{ij}X^iX^j dU^2$ in (18), traded for "time"-dependent" transverse metric  $a_{ij}(u)$  (while leaving U = uunchanged).

## EXAMPLES

#### 0. Restriction of flat Minkowski space

 $dr^2 + 2dt ds$ 

to t = 0 is Carroll manifold, upon which restriction e = 0 of Bargmann group acts consistently with Carroll action.



Linearly polarized "sudden burst"  $\sim$  Gaussian profile ( $\sim$  anisotropic oscillator with time-dependent frequency)

$$K_{ij}(u)X^{i}X^{j} = \frac{e^{-u^{2}}}{\sqrt{\pi}} ((X^{1})^{2} - (X^{2})^{2}).$$
 (29)

Fig.8 "Time" evolution of wave for "sudden burst" with Gaussian profile  $\mathcal{A}(u) = \exp[-u^2]$ .

Sandwich wave:  $K(u) \neq 0$  only in "wave zone"  $U_i < U < U_f$ . Assumption : metric Minkowski in "before-zone"  $U < U_i$  and flat in "afterzone"  $U_f < U$ .

$$\mathcal{A}(U) = 2k\,\delta(U) \tag{30}$$

 $k \in \mathbb{R}$ . Wave zone suppressed,  $U_i = U_f = 0$ . SL eqn. (27) solved by

$$P(u) = 1 + u \theta(u) c_0 \tag{31}$$

where  $\theta(u)$  Heaviside,  $c_0 = \frac{1}{2}\dot{a}(0+)$  initial "speed" of transverse metric. Can be chosen  $c_0 = k \operatorname{diag}(1, -1)$ .

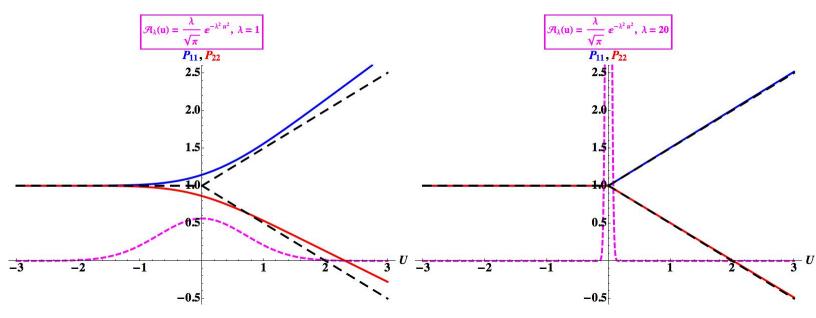


Fig.9. Numerical solution of S-L eqn (27) for profile  $\mathcal{A}_{\lambda}(U) = (\lambda/\sqrt{\pi}) e^{-\lambda^2 U^2}$  shows that components of diagonal matrix  $P_{\lambda}(U)$  approach, for large  $\lambda$ , those of impulsive wave [in **dashed black**].

$$a(u) = \begin{cases} 1 & \text{for } u \le 0, \\ (1+uc_0)^2 & \text{for } u \ge 0. \end{cases}$$
(32)

More generally

$$\mathcal{A}_{\lambda}(U) = \frac{\lambda}{\sqrt{\pi}} e^{-\lambda^2 U^2}.$$
 (33)

Squeezing Gaussians to Dirac  $\delta$  by letting  $\lambda \to \infty$ , components of  $P_{\lambda}(u)$  and of transverse metric  $a_{\lambda}(u) = P_{\lambda}^{T}(u)P_{\lambda}(u)$  tend to those of impulsive wave.

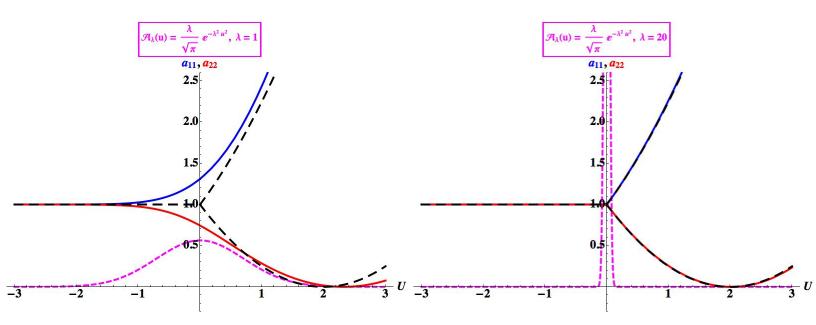


Fig.10. Squeezing Gaussians  $\mathcal{A}_{\lambda}$  to Dirac  $\delta$ , transverse metrics  $a_{\lambda}(u)$  (in red and blue) tend to that of impulsive wave in BJR coordinates, depicted in dashed black lines.

#### Carroll boost for impulsive GW

Boost implemented as  $x \to x + H(u)\mathbf{b}, v \to v - \mathbf{b} \cdot x - \frac{1}{2}\mathbf{b} \cdot H(u)\mathbf{b}$  cf. (23). For impulsive

$$H(u) = u P^{-1}(u)$$
(34)  
$$P = \begin{cases} 1 & u \le 0 \\ diag(1+u/2, 1-u/2) & u \ge 0 \end{cases}$$
(35)

$$H = \text{diag}(H_{+}, H_{-}) = \begin{pmatrix} \frac{u}{1+u/2} & \\ & \frac{u}{1-u/2} \end{pmatrix} \ u \ge 0.$$
(36)

Boost with  $\mathbf{b} = (b_+, b_-)$  implemented as,

$$x_1 \to x_1 + \frac{u}{1 + u/2} \frac{b_+}{1 + u/2}$$
 (37a)

$$x_2 \to x_2 + \frac{u}{1 - u/2} b_-$$
 (37b)

$$v \to v - (x_1 b_+ + x_2 b_-) - \frac{1}{2} \left( \frac{u}{1 + u/2} b_+^2 + \frac{u}{1 - u/2} b_-^2 \right)$$
 (37c)

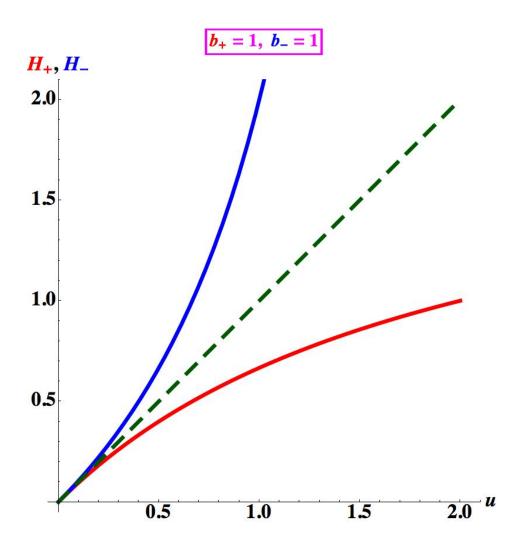


Fig.11 Boost acts on impulsive space-time according to  $x \to x + H\mathbf{b}, v \to v - \mathbf{b} \cdot x - \frac{1}{2}\mathbf{b} \cdot H(u)\mathbf{b}$ .  $H = \text{diag}(H_+, H_-)$ but components differ considerably from usual Galilei implementation  $H_{Gal} = u \, \text{Id}$ .

### Polarized oscillating GW with Gaussian envelope

In Brinkmann (B) coordinates (X, U, V) profile of plane GW given by symmetric & traceless  $2 \times 2$  matrix  $H(U) = H_{ij}(U)$ 

$$\delta_{ij} dX^{i} dX^{j} + 2dUdV + H_{ij}(U)X^{i}X^{j}dU^{2}, \quad (38a)$$
$$H_{ij}(U)X^{i}X^{j} = \qquad (38b)$$
$$\frac{1}{2}\mathcal{A}(U)\left((X^{1})^{2} - (X^{2})^{2}\right) + \mathcal{B}(U)X^{1}X^{2}$$

 $\mathcal{A} \& \mathcal{B}(U)$  amplitudes of + and  $\times$  polarization states.

(geodesic) eqns motion

$$\frac{d^2 \mathbf{X}}{dU^2} - H(U) \mathbf{X} = 0, \quad H(U) = \frac{1}{2} \begin{pmatrix} \mathcal{A} & \mathcal{B}(U) \\ \mathcal{B}(U) & -\mathcal{A} \\ (39a) \end{pmatrix},$$
(39a)

$$\frac{d^{2}V}{dU^{2}} + \frac{1}{4}\frac{d\mathcal{A}}{dU}\left((X^{1})^{2} - (X^{2})^{2}\right) + \mathcal{A}\left(X^{1}\frac{dX^{1}}{dU} - X^{2}\frac{dX^{2}}{dU}\right) + \frac{1}{2}\frac{d\mathcal{B}}{dU}X^{1}X^{2} + \mathcal{B}\left(X^{2}\frac{dX^{1}}{dU} + X^{1}\frac{dX^{2}}{dU}\right) = 0.$$
(39b)

#### **Circularly Polarized Waves**

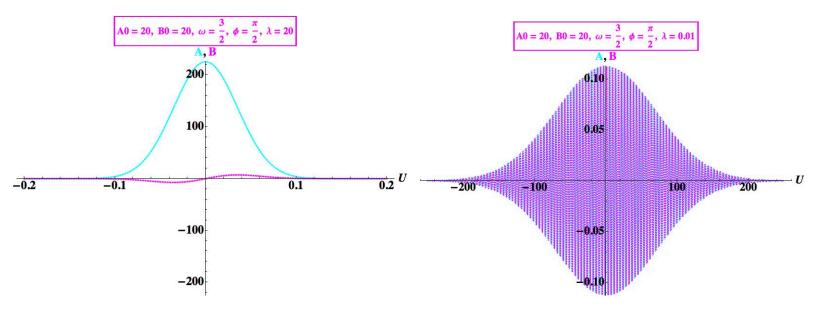
• If  $\mathcal{A}(U) = 0$  or  $\mathcal{B}(U) = 0$ , wave linearly polarized.

Polarized waves approximating the sandwich by Gaussian,

$$\mathcal{A}(U) = A_0 \frac{\lambda}{\sqrt{\pi}} e^{-\lambda^2 U^2} \cos(\omega U), \qquad (40a)$$

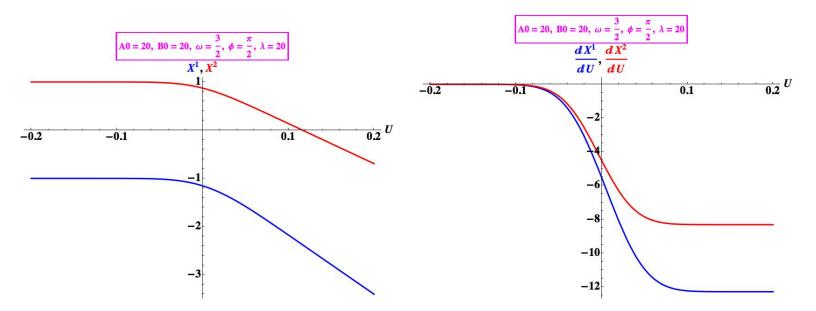
$$\mathcal{B}(U) = B_0 \frac{\lambda}{\sqrt{\pi}} e^{-\lambda^2 U^2} \sin(\omega U).$$
 (40b)

choose  $A_0 = B_0$ .



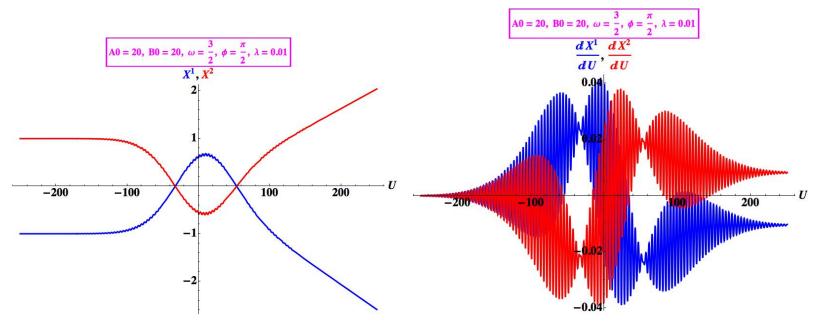
Profile of circularly polarized sandwich waves for large / small  $\lambda$ . For  $\lambda \to \infty$  Gaussian profile approximates polarized impulsive wave; for  $\lambda \to 0$  becomes weak but wide. (Note different scales).

Numerically calculated trajectories and velocities hint at following behavior: • In large- $\lambda$  regime Gaussian thin & high, motion roughly along straight lines with constant velocity, except in short ( $\sim O(1/\lambda)$  transitory "inside" region, where trajectory sharply bent, velocity changes rapidly from zero to non-zero value — reminiscent of motion in impulsive wave.



Trajectories / velocities after passage of circularly polarized Gaussian sandwich wave (40) in impulsive limit  $\lambda \to \infty$ . blue and red colors refer to transverse components  $X^1$  and  $X^2$ . Trajectory bent in (narrow) inside zone & then escapes with non-vanishing constant velocity in flat after-zone.

• In <u>small- $\lambda$ </u> regime profile wide & low, with many oscillations in inside zone. Apart of fine "denting", motion "reasonably regular" in inside zone. Effect of "denting" important when velocity is plotted. Trajectories suffer weak rotation.



Trajectories / velocities for small  $\lambda$ , describing particle motion in wide but weak circularly polarized Gaussian sandwich GW (40). blue and red colors refer to transverse components  $X^1$  and  $X^2$ . Particle initially at rest has complicated motion in inside zone however escapes with non-zero constant velocity in flat after-zone. In outside zones,  $U < U_i \& U > U_f$ , everything smoothes out for both regimes: the trajectory (just like on the linearly polarized case) follows straight lines with constant velocity.

Large and small- $\lambda$  regimes differ in inside zone: motion in after zone always simple.

### boost implemented

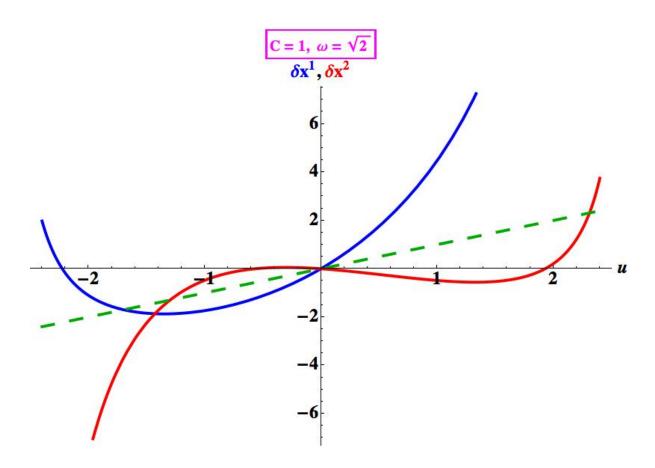


Fig.12 Boost acts on polarized oscillating GW with Gaussian envelope according to

$$x \rightarrow x + H\mathbf{b}, v \rightarrow v - \mathbf{b} \cdot x - \frac{1}{2}\mathbf{b} \cdot H(u)\mathbf{b}$$

Components differ considerably from Galilei implementation H = u Id.