Schrödinger–Newton Equation: Four Fundamental Catches and Attempts to Relax Them

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In 1984, Schrödinger-Newton equation opened the door to a non-relativistic regime of nano/micro-quantum-mechanics, instead of quantum cosmology, where quantum and gravity are equally important. The fundamental difficulties, well-known and less known ones, of this non-linear equation are summarized. Some concepts to relax or merely cope with them are interpreted. of quantum mechanics and special relativity (Shimony)

- apparent action-at-a-distance in EPR situation
- non-locality in Bell formulation

Still:

action-at-a-distance (AAD) & faster-then-light (FTL) communication remain impossible.

Reason: linear structure of quantum mechanics Non-linear modifications open door to **FTL** communication! (Gisin)

$$i\hbar\frac{d\psi}{dt} = \hat{H}\psi + \hat{V}_{\psi}\psi$$

allows for **FTL** communication for whatever small (non-trivial) \hat{V}_{Ψ} .



Semiclassical Gravity in Cosmology

Effective theory in cosmology/astrophyics: coupling quantized matter and classical gravity, **Semiclassical Einstein Equation**:

$$G_{ab} = rac{8\pi G}{c^4} \left\langle \hat{T}_{ab}
ight
angle_\psi$$

Four fundamental catches:

- Fake action-at-a-distance AAD
- Faster-then-light FTL communication
- No local autonomy NoLA
- No Born statistical interpretation NoBorn

Unrelated to relativity, gravity. Related to quantum nonlinearity induced by $\langle \hat{T}_{ab} \rangle_{\psi}$. Transparent in the nonrelativistic limit.

Semiclassical Gravity in the Lab

Newtonian limit of Semiclassical Einstein Equation

$$G_{00} = 8\pi G c^{-4} \left\langle \hat{T}_{00} \right\rangle_{\psi}$$

where $G_{00} = 2c^{-2}\Delta\Phi$, $\hat{T}_{00} = \hat{\varrho}c^2$, hence c cancels:
 $\Delta\Phi = 4\pi G \langle \hat{\varrho} \rangle_{\psi}.$
 $\Phi_{\psi}(x) = -G \int \frac{d^3r}{|x-r|} \langle \hat{\varrho}(r) \rangle_{\psi}.$

The Schrödinger-Newton Equation (SNE):

$$i\hbar \frac{d\psi}{dt} = \hat{H}\psi + \int \Phi_{\psi}(r)\hat{\varrho}(r)d^{3}r\psi$$

Effective theory? Might be fundamental! (D. 1984, Penrose 1996) All four catches survive: **AAD, FTL, NoLA, NoBorn**

... for c.o.m. free motion of "pointlike" big mass M:

$$\Phi_{\psi}(x) = -GM \int \frac{|\psi(r)|^2}{|x-r|} d^3r$$

$$i\hbarrac{d\psi(x)}{dt}=-rac{1}{2M\hbar^2}
abla^2\psi(x)+M\Phi_\psi(x)\psi(x)$$

- Self-field Φ_ψ means self-attraction
- Self-attraction yields solitons, size $\sim (\hbar^2/\textit{GM}^3)$
- Testable soon in nano-Quantum-Mechanics:

"Quantum Gravity in the Lab"

- One-soliton: stable ground state. Might be the natural localized state of macroobjects.
 Soliton size: astronomic for elementary particles, subsubmicroscopic for large objects, crossover at 10⁻⁵cm.
- Two-soliton: total mass *M* is shared between the parts. They **attract each other** gravitationally. Can penetrate each other. Any classical two-body motion is stable approximately, e.g. Keplerian motion.

Example: two solitons, mass M/2 each, can oscillate along a line, penetrating each other, with the classical period $T = \pi \sqrt{\ell^3/2GM}$ where ℓ is the amplitude of their distance.

• . . .

Alice and Bob, far away from each other, own a qubit and a SN-Kepler Cat, respectively, in entangled state:

$$\frac{1}{\sqrt{2}} \Big(|\!\uparrow\rangle \otimes | \text{``half-cat''-1} \rangle + |\!\downarrow\rangle \otimes | \text{``half-cat''-2} \rangle \Big)$$

At t = 0 Alice i) does nothing or ii) measures her qubit. At some time t > 0 Bob measures position of his Cat. Bob finds that his Cat is on the orbit [case i)] or that the Cat left the orbit [case ii)].

- AAD: Alice local action changed dynamics at Bob's location.
- FTL: She did it superluminally if their distance was large.
- **NoLA**: Bob can in general not predict the dynamics of his system.
- NoBorn: ...

Nonlinear SNE without FTL & NoBorn ?

Delay collapse (caused by Alice) causally at Bob's location! (Kent 2005, Bedingham 2016, Helou & Chen 2017) Standard collapse is instantaneous:



Proposal: Bob's local state be not collapsed by Alice because her measurement falls outside Bob's past-lightcone.

Stochastic Semiclassical Gravity: AAD, FTL, NoLA, NoBorn Have Gone!

Persecute Schrödinger–Newton Cats! (Tilloy & D. 2016-17) Assume universal (spontaneous) measurement of mass distribution $\hat{\varrho}(x)$, yielding the measured outcome $\varrho(x, t)$ of the form

$$\varrho = \langle \hat{\varrho} \rangle_{\psi} + \mathsf{noise}$$

On r.h.s. of semi-classical Newton-Poisson eq., replace $\langle \hat{\varrho} \rangle_{\psi}$ by ϱ :

$$\Delta \Phi_{\psi} = 4\pi G \Big(\langle \hat{\varrho} \rangle_{\psi} + \text{noise} \Big)$$

Optimize trade-off between noise and precision of $\hat{\varrho}$ -measurement. Result: the **Gravity-Related Spontaneous Collapse** theory, known from alternative considerations (Penrose & D.)

Statistics Require Linearity

- Suppose any **dynamics** $\hat{
 ho}^f = \mathcal{M}[\hat{
 ho}^i]$, not necessarily linear.
- Consider statistical **mixing** of $\hat{\rho}_1, \hat{\rho}_2$ with weights $\lambda_1 + \lambda_2 = 1$:

$$\hat{\rho} = \lambda_1 \hat{\rho}_1 + \lambda_2 \hat{\rho}_2$$

In von Neumann standard theory mixing and dynamics are interchangeable:

$$\mathcal{M}[\lambda_1 \hat{\rho}_1 + \lambda_2 \hat{\rho}_2] = \lambda_1 \mathcal{M}[\hat{\rho}_1] + \lambda_2 \mathcal{M}[\hat{\rho}_2]$$

Recognize the condition of $\mathcal{M}\mbox{'s linearity!}$

- Interchangeability excludes deterministic non-linear Schrödinger equations
- Without interchangeability statistical interpretation collapses

Catch NoBorn is non-quantum, it's classical statistical! (D.: *A Short Course in Quantum Information Theory*, Springer, 2007, 2011)

However hard the **conflict between non-linearity and statistics** is, Schrödinger–Newton equation deserves attention in foundations. Just we should keep in mind catches: The SNF

- allows for
 - fake action-at-a-distance (maybe extreme weak)
 - faster-than-light communication (maybe too hard to realize)
- does not allow for
 - local autonomous dynamics (unless you have a pure state)
 - Born statistical interpretation (maybe a substitute works?)