

# Trans-Planckian non-relativistic motion is not an oxymoron

Lajos Diósi

Wigner Center, Budapest

11 June 2019, Vaxjø



Supported by

Foundational Questions Institute & Fetzer Franklin Fund, National  
Research Development and Innovation Office of Hungary

Grant number K12435

# No QG effects below $10^{19}$ GeV?

length	$\ell_{\text{Pl}} = \sqrt{\frac{\hbar G}{c^3}}$	$= (1.6 \times)$	$10^{-33}$	cm
time	$\tau_{\text{Pl}} = \sqrt{\frac{\hbar G}{c^5}}$	$= (5.4 \times)$	$10^{-44}$	s
mass	$m_{\text{Pl}} = \sqrt{\frac{\hbar c}{G}}$	$= (2.2 \times)$	$10^{-5}$	g
temperature	$T_{\text{Pl}} = \sqrt{\frac{\hbar c^5}{G k_B^2}}$	$= (1.4 \times)$	$10^{32}$	K
energy	$E_{\text{Pl}} = \sqrt{\frac{\hbar c^5}{G}}$	$= (1.4 \times)$	$10^{19}$	GeV
		$= (3.0 \times)$	$10^{16}$	erg

High Energy Physics wisdom:

- QG is believed to dominate on the trans-Planckian scale
- QG effects are totally ingorable below the Planck-scale
- Non-relativistic (NR) QM can not overlap with QG
- Planckian or trans-Planckian slow  $\ll c$  motion is an oxymoron
- ... like “QG in the Lab”

# Still there is a loophole to QG in the lab

Google's wisdom ("quantum gravity in the lab"):

- 10 results before 2000
- 12.600 results until 2019 (7 June)

How can this deal exist?

My wisdom hint:

- HEP requests  $10^{19} \text{ GeV}$
- ... per elementary particle — unavailable after Big Bang
- $10^{19} \text{ GeV} \sim 10^{16} \text{ erg}$  per massive object — available in the lab

$$E = 10 \text{ kg} \times (10 \text{ km/s})^2 = 10^{16} \text{ erg} \sim E_{\text{Pl}}$$

Loophole from high energies toward the lab: massive d.o.f. to test Planck-scale!

Next: Planck length  $\ell_{\text{Pl}}$  can be hit earlier than Planck energy  $E_{\text{Pl}}$ .

# Trans-Planckian de Broglie wavelength

$$\underbrace{\lambda_{\text{deBroglie}}}_{\frac{2\pi\hbar}{mv} \sqrt{1-v^2/c^2}} \sim \underbrace{\ell_{\text{Planck}}}_{\sqrt{\frac{\hbar G}{c^3}}}$$

$\lambda_{\text{dB}}$  sinks to and below  $\ell_{\text{Pl}}$  in two ways:

- Relativistic way: elementary particles velocity  $v$  closes  $c$ , requests  $10^{19}\text{GeV}/\text{particle}$  — forget it!
- NR way: velocity  $v$  remains  $\ll c$  but *mass grows macroscopic* — might be viable in the lab

$$\lambda_{\text{dB}} = \frac{2\pi\hbar}{\mathbf{10g} \times \mathbf{10km/s}} \sim \mathbf{10^{-33}cm} \sim \ell_{\text{Pl}}$$

1) NR (trans-)Planckian  $m = 10\text{g}$  Schrödinger cat:

$$|v = -10\text{km/s}\rangle + |v = +10\text{km/s}\rangle$$

2) NR (trans-)Planckian vibration of  $m = 10\text{kg}$  body:

$$\lambda_{\text{dB}} = \frac{2\pi\hbar}{m\omega a} = \frac{2\pi\hbar}{10\text{kg} \times 100\text{kHz} \times 10^{-2}\text{cm}} \sim 10^{-33}\text{cm} \sim \ell_{\text{Pl}}$$

# What happens at the Planck scale?

- Bronstein 1935: Notion of space-time continuum may be lost.
- Wheeler 1962: Foamy space-time
- Hawking 1983: Non-unitary (decohering) dynamics

No accepted dynamics for scales  $\ell_{\text{Pl}}$  (Garay 1995, Hossenfelder 2013)

Phenomenology:  $\ell_{\text{Pl}}$ -scale metric fluctuations

Simplest choice: stochastic conformal fluctuations

$$\{g_{ab}\} = (1 + h)\text{diag}(1, -1 - 1 - 1), \quad |h| \ll 1$$

Lorentz-invariant correlation (D 2019):

$$\mathbb{E}h(x)h(y) = \ell_{\text{Pl}}^2 \int \frac{d^4k}{(2\pi)^4} \frac{\theta(-k^2)}{-k^2} e^{-ik(x-y)}, \quad (1)$$

Write  $h$  as  $h = 2\Phi/c^2$ , take  $c \rightarrow \infty$ , then  $c$  cancels!

$$\mathbb{E}\Phi(t, \mathbf{x})\Phi(s, \mathbf{y}) = \frac{\hbar G}{|\mathbf{x} - \mathbf{y}|} \delta(t - s).$$

# Trans-Planckian NR Schrödinger cats decohere

... and die.

NR QM becomes modified by the stochastic Newton-potential:

$$\frac{d\Psi}{dt} = -\frac{i}{\hbar} \left( \hat{H} + \sum m_n \Phi(\hat{x}_n, t) \right) \Psi$$

$$\mathbb{E}\Phi(t, \mathbf{x})\Phi(s, \mathbf{y}) = \frac{\hbar G}{|\mathbf{x} - \mathbf{y}|} \delta(t - s),$$

causing G-related decoherence in NR QM (D 1987, Penrose 1996).

Trans-Planckian NR Schrödinger cat decays

$$\frac{|-10\text{km/s}\rangle + |10\text{km/s}\rangle}{\sqrt{2}} \Rightarrow \begin{cases} |-10\text{km/s}\rangle & 50\% \\ |10\text{km/s}\rangle & 50\% \end{cases}$$

at the DP-decay time

$$10^{-15} - 10^{-9} \text{s}.$$

# Summary

- NR wavefunction of (isolated) massive d.o.f. acquires trans-Planckian structure.
- Hence NR QM should get some corrections for large masses.
  - We have no theory there — since we have no safe QG theory.
  - Yet, QG-phenomenology says: metric fluctuates heavily there.
  - Fluctuations are ignorable for any NR atomic object.
  - Cumulative effect of fluctuations does matter for massive d.o.f.
  - Their primary effect is decoherence of massive d.o.f.
  - DP version seems to clear NR states of their trans-Planckianity.
- **Trans-Planckian & non-relativistic? Yes, in massive d.o.f.,** theorists should do something about QM there.
- **QG & lab? Yes! Quantum control of "big-enough" d.o.f.** would come soon (10yy? — ask e.g. Aspelmayer).