## The hadronization line in stringy quark matter



#### Dictionary.refrence.com/search?q=stringy

- <u>Modern Language Assciation,</u> <u>Random House Inc. 2008</u>
- Resembling a string or strings, consisiting of strings or stringlike pieces: stringy weeds; a stringy fiber
- 2. Coarsely or toughly fibrous, as meat
- 3. Sinewy or wiry, as a person
- 4. Ropy, as a glutinous liquid

- <u>The American Heritage</u> <u>Dictionary of the English</u> <u>Language, Houghton Mifflin Co.</u> <u>2006</u>
- 1. Consisting of, resembling or containing strings or a string
- 2. Slender and sinewy; wiry
- 3. Forming strings, as a viscous liquid; ropy.

#### Dictionary.refrence.com/search?q=stringy

- <u>On-line Medical Dictionary, 1997-98 Academic Medical Publishing &</u> <u>CancerWEB</u>
- Consisting of strings or small thraeds; fibrous; filamentous; as a stringy root.
- 2. Capable of being drawn into a string; as a glutinous substance; ropy; viscid; gluely.
- 3. Stringy bark, a name given in Australia to several trees of the genus Eucalyptus (as E. amygdalina, obliqua, capitellata, macrorhyncha, piperita, pilularis and tetradonta), which have a fibrous bark used by the aborigines for making cordage and cloth.

The hadronization line in stringy quark matter

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- How can be E / N = 6 T ?
- Stringy corrections to QGP
- High-T equation of state
- The zero pressure line

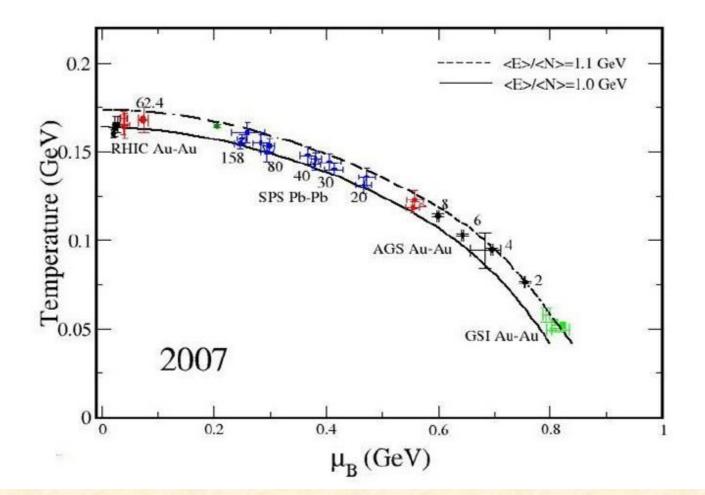
NFQCD Workshop, Yukawa Institute, Kyoto, February 13, 2008.



## **Statistical Model**

#### Statistical model

Talk: Cleymans



## T = 167 MeVE / N = 1 GeV



Statistical Model: hadronization point around  $\mu = 0$  (RHIC, LHC)

## m = 750 MeV

#### Massive hadrons (rho?)

## E/N = 3T

## m = 0

#### Ideal gas of radiation

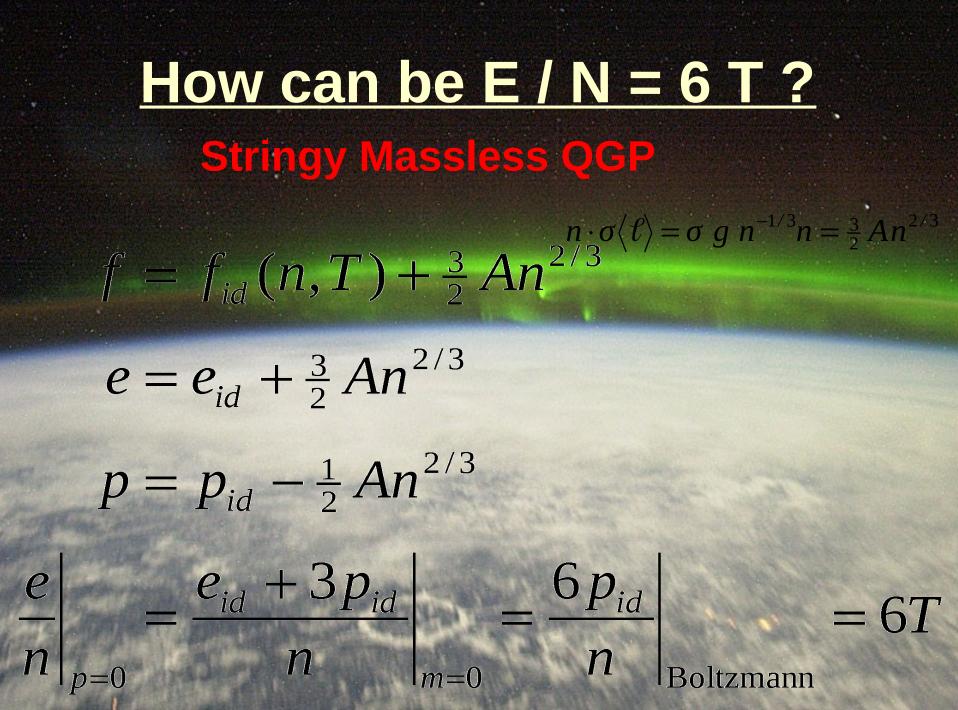
# $e = \sigma T^{4} + B$ $p = \frac{1}{3}\sigma T^{4} - B \ge 0$ $m \approx \frac{1}{3}\sigma T^{3}$

 $E/N = e/n \leq 41$ 

Bag Model for QGP

## How can be E / N = 6 T ?Stringy Massless QGP $f = f_{id}(n,T) + \frac{3}{2}An^{2/3}$ $e = e_{id} + \frac{3}{2}An^{2/3}$ $p = p_{id} - \frac{1}{2}An^{2/3}$ $e_{id} + 3p_{id}$ $6p_{id}$ = 6Tn n m=0

e



## How can be E/N = 6T?

This is more generally true!

correction depends on



 $p_{id} = T \cdot \sum n_i$ 

- constituents are massless  $p_{id} = T^4 \cdot \phi \left( \frac{\mu_B}{T} \right)$
- Boltzmann approximation is acceptable

$$C = \sum_{i} C_{i} n_{i}$$
$$f = f_{id} + \frac{1}{1-\gamma} A C^{1-\gamma}$$

color density

free energy density fractional power

chemical potential

$$p = p_{id} - \frac{\gamma}{1-\gamma} A c^{1-\gamma}$$

 $\mu_i = \mu_{id} + Ac^{-\gamma}c_i = q_i \mu_B$ 

$$e = e_{id} + \frac{1}{1-\gamma} A c^{1-\gamma}$$

pressure

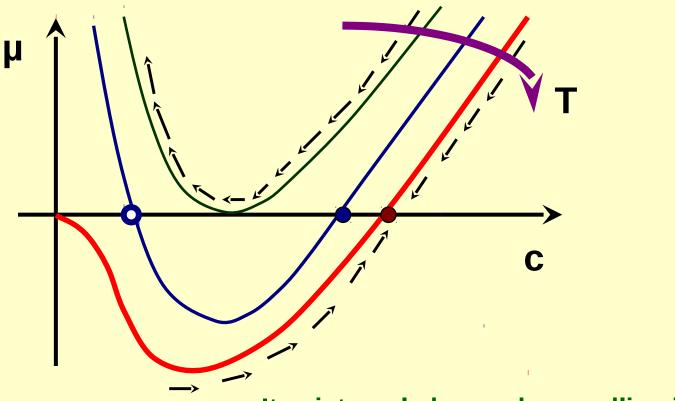
energy density

**Consistent solution for color density in equilibrium:** 

$$c = \sum_{i} c_{i} d_{i} n_{id} (T, q_{i} \mu_{B} - Ac_{i} c^{-\gamma})$$
$$\implies c = c (T, \mu_{B})$$

It exists only beyond an endline in T -  $\mu$ 

**Consistent solution for color density in equilibrium:** 



It exists only beyond an endline in T -  $\mu$ 

**One massless component, Boltzmann approximation:** 

$$c = \frac{d}{\pi^2} T^3 e^{q\mu_B/T} e^{-Ac^{-\gamma}/T}$$

$$z = \frac{A}{T}c^{-\gamma}, \quad \alpha = \frac{A}{T} \left(\frac{d}{\pi^2}T^3 e^{q\mu_B/T}\right)^{-\gamma}$$

$$z = \alpha e^{\gamma z}, \quad z = -\frac{1}{\gamma}W(-\gamma \alpha)$$

The solution is related to Lambert's function

**Endline: last possible solution** 

Boltzmann approximation

$$z = \alpha e^{\gamma z}, \quad 1 = \gamma \alpha e^{\gamma z}$$

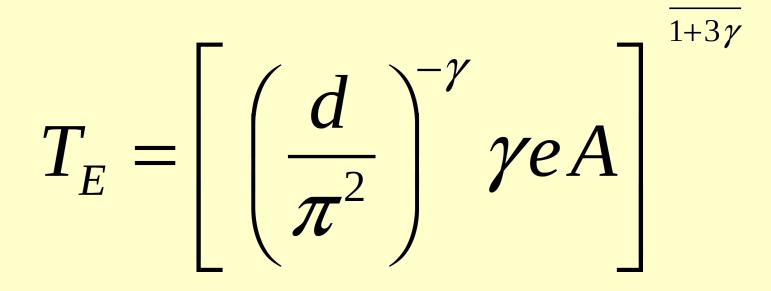
$$\Rightarrow Z_E = \frac{1}{\gamma} \Rightarrow \alpha_E = \frac{1}{\gamma} e^{-1}$$

$$\mu_{B} = \frac{3\gamma + 1}{\gamma q} T \ln \frac{T_{E}}{T}$$

It is near to the zero pressure line for low chemical potential

Boltzmann approximation

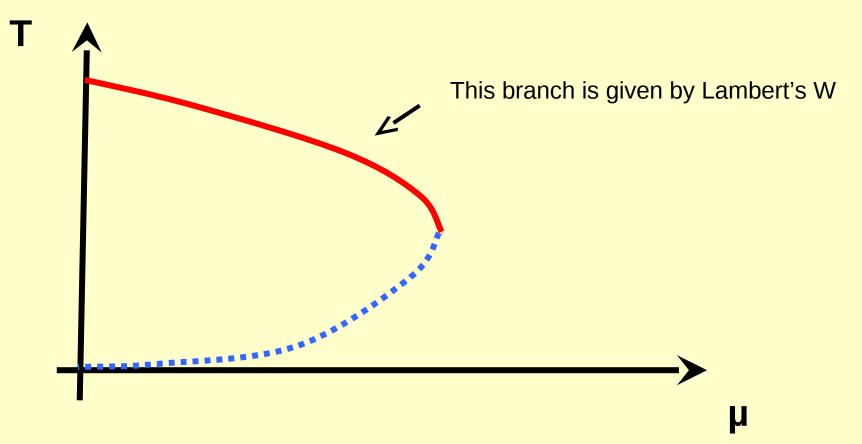
End temperature: last possible solution at  $\mu = 0$ 



It is near to the zero pressure line for low chemical potential

**Endline: last possible solution** 





## The zero pressure line

Zero pressure in the Boltzmann approximation

$$p = cT\left(1 - \frac{\gamma}{1 - \gamma}z\right)$$

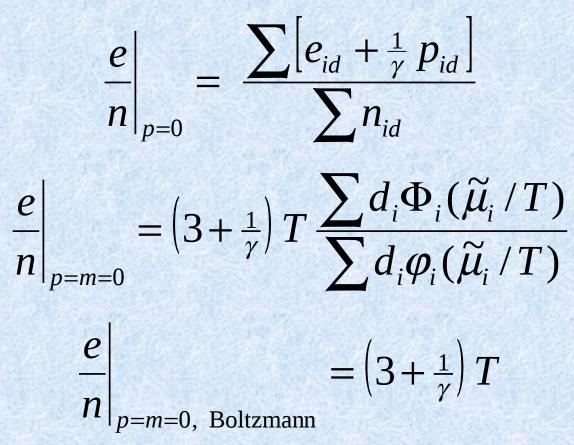
$$\Rightarrow z_0 = \frac{1-\gamma}{\gamma} \Rightarrow \alpha_0 = \left(\frac{1}{\gamma} - 1\right) e^{\gamma - 1}$$

$$\mu_{B} = \frac{3\gamma + 1}{\gamma q} T \ln \frac{T_{0}}{T}, \quad \frac{T_{0}}{T_{E}} = (1 - \gamma)^{-\frac{1}{3\gamma + 1}} e^{-\frac{\gamma}{3\gamma + 1}}$$

Endline and zero pressure line are relatively close! (T0  $\approx$  1.04 TE)

## **E/N at the zero pressure line**

#### **General relations**



This is independent of the value of the string tension!

## <u>High – T equation of state</u>

$$\frac{e - 3p}{T^4} = \frac{1 + 3\gamma}{1 - \gamma} \frac{A}{T^4} c^{1 - \gamma} \sim T^{-3\gamma - 1}$$

lattice QCD :

$$3\gamma + 1 = 2 \rightarrow \gamma = \frac{1}{3}$$

## **High-T behavior of lattice eos**

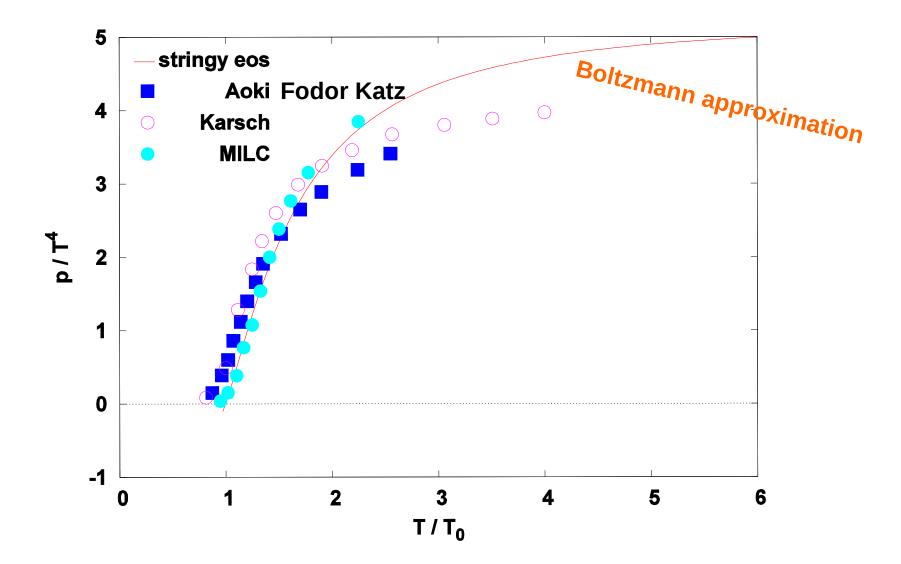
#### Interaction measure log-log plot 100 Karsch et.al. ⊿ $m_0$ exp(-2\*log(x)+1.23)10 $T^{2}$ 0.1 Δ 0.01 10

(e-3p) / T<sup>4</sup>

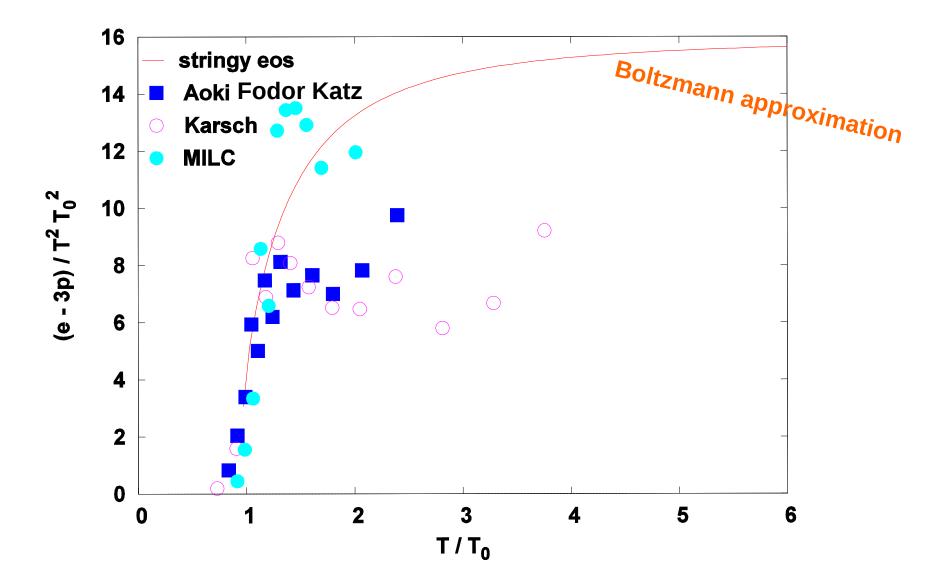
0.1

T / T<sub>c</sub>

## **High-T behavior of lattice eos**



## **High-T behavior of lattice eos**



### Where to stop with the

## **Boltzmann approximation?**

Fermi integral  $\approx$  Boltzmann integral for negative or zero  $\mu$ .

 $q\mu_{\rm B} - Ac^{-\gamma} \le 0$ 

# $\frac{q\mu_B}{T} = \left(3 + \frac{1}{\gamma}\right) \ln \frac{T_E}{T} \le z_E = \frac{1}{\gamma}$ $T_E > T > T_E e^{-1/2} \approx 0.6 T_E$

The endline does not turn back!

 $T_{\rm max} = T_E \ e^{-1}$ 

## The zero pressure line

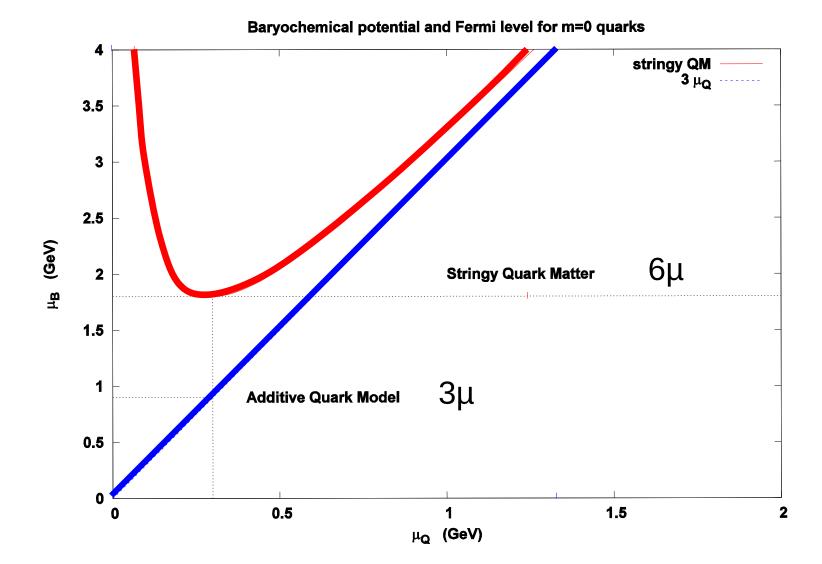
Fermions (quarks only) at T = 0

$$c = \frac{dc_q}{6\pi^2} (q\mu_B - Ac^{-1/3}c_q)^3$$

$$c = c_q \left[ \frac{q\mu_B}{2B} + \sqrt{\frac{q^2\mu_B^2}{4B^2} - \frac{\tilde{A}}{B}} \right]^3, \quad \frac{e}{n} \Big|_{p=0, T=0} = q\mu_B$$

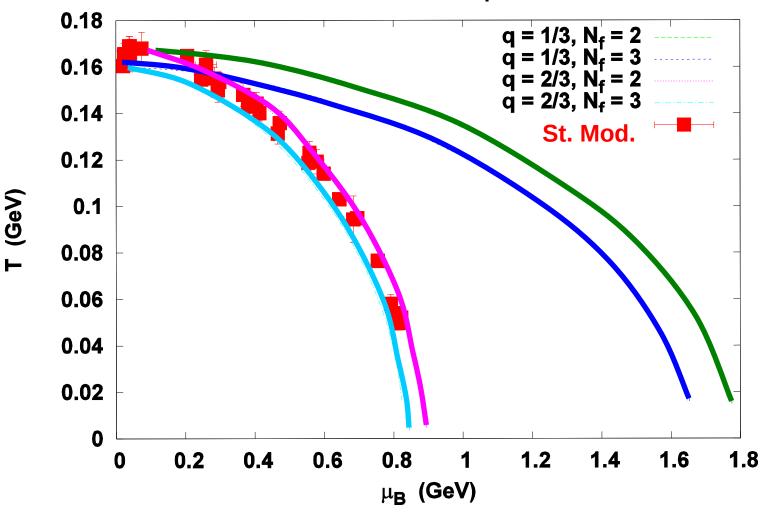
$$\mu_0^Q = \frac{3}{2\sqrt{2}} \mu_E \approx 1.06\mu_E \approx 10.4 T_0^{QGP} \approx 1.74 \text{ GeV}$$

## **The endpoint at T=0**



## The zero pressure line

Stringy massless QGP p = 0 line ( $c_i = 1$ , A = 0.04564 GeV<sup>2</sup>)



## Summary

- E / N = 6T only from stringy massless QGP
- fractional power is 1/3 from high-T lattice eos
- zero pressure line and endpoint of solution are close
- at finite baryon density and zero temperature the hadronization line is at mu\_B = 3 k\_F
- at any temperature IR effects may be present (already a small string tension makes E / N = 6T)

If the accumulation of false beliefs is cleared away, Enlightenment will appear. But, strange enough, when people attain Enlightenment, they will realize that Without false beliefs there could be no Enlightement.

(The Teaching of Buddha)

## Bibliography

• **arXiv: 0801.3998**, T.S.Biro and J. Cleymans: The hadronization

line in stringy matter

- hep-ph / 98083941, T.S.Biro, P.Levai, J.Zimanyi and C.Traxler: Hadronization in heavy ion collisions, PRC59, ....., 1999
- T.S.Biro, A.A.Shanenko and V.D.Toneev: Toward Thermodynamic
   Consistency of Quasiparticle Picture, Phys.Atomic Nuclei 66, 982, 2003
- J.Cleymans, H.Oeschler, K.Redlich, S.Wheaton, PRC73, 034905, 2006.



## FAQ

- What is actually non-perturbative at T > Tc ? (low Q2 pairs)
- What is non-particle like (stringy) at T > Tc ? (interaction)
- Are all color charges stringed at T > Tc? No! (a small fraction suffices for the effect.)
- Asymptotic freedom is incomplete at any finite temperature. Why and what is the quantitative measure of this effect?  $\rightarrow$

## Arguments

Non-perturbative effects at arbitrary high temperature:

- **1. Thermal distribution of Q<sup>2</sup>**
- 2. NP order parameter: cut-off in Q<sup>2</sup>
- 3. Its thermal expectation value  $\rightarrow$  order of NP effects
- 4. high-T expansion
- □□→ high-T NP terms in EoS (pressure, int.mesure)

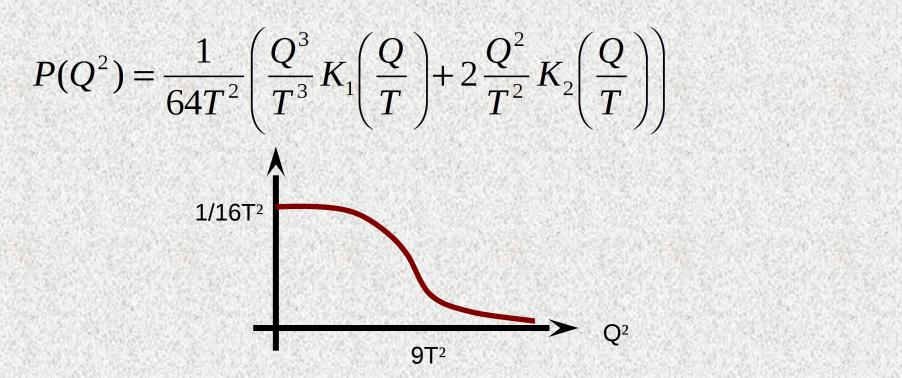
## **Pressure: NP effects at any T**

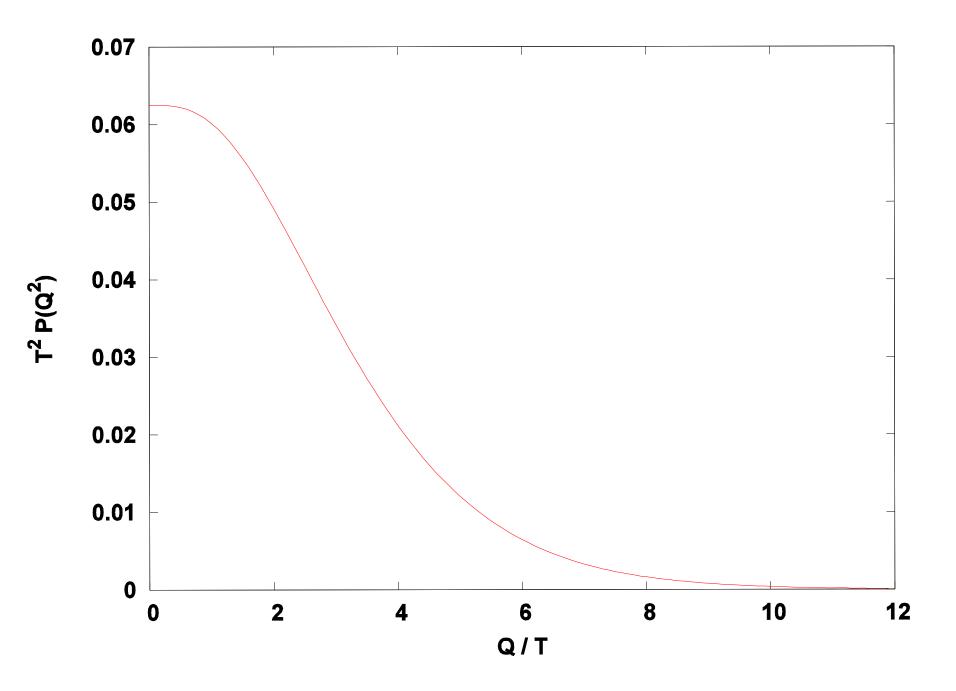
$$p = p_{p} \int_{\Lambda^{2}}^{\infty} P(Q^{2}) dQ^{2} + p_{NP} \int_{0}^{\Lambda^{2}} P(Q^{2}) dQ^{2}$$
$$\int_{0}^{\Lambda^{2}} P(Q^{2}) dQ^{2} = \int_{0}^{\Lambda^{2}/T^{2}} f(x) dx$$
$$p = p_{p} - (p_{p} - p_{NP}) \begin{cases} 1 & (T << \Lambda) \\ \Lambda^{2}/T^{2} & f(0) & (T >> \Lambda) \end{cases}$$

If it were f(0) = 0, then the QGP pressure would be free of NP effects!

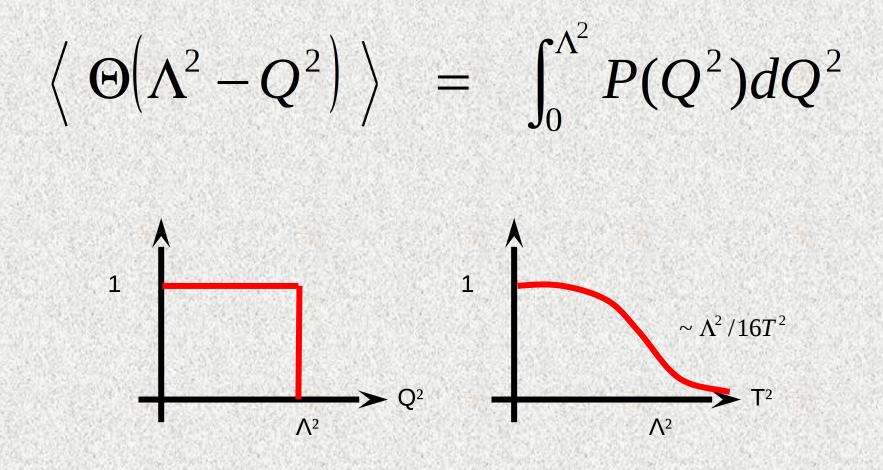
## **Thermal distribution of Q**<sup>2</sup>

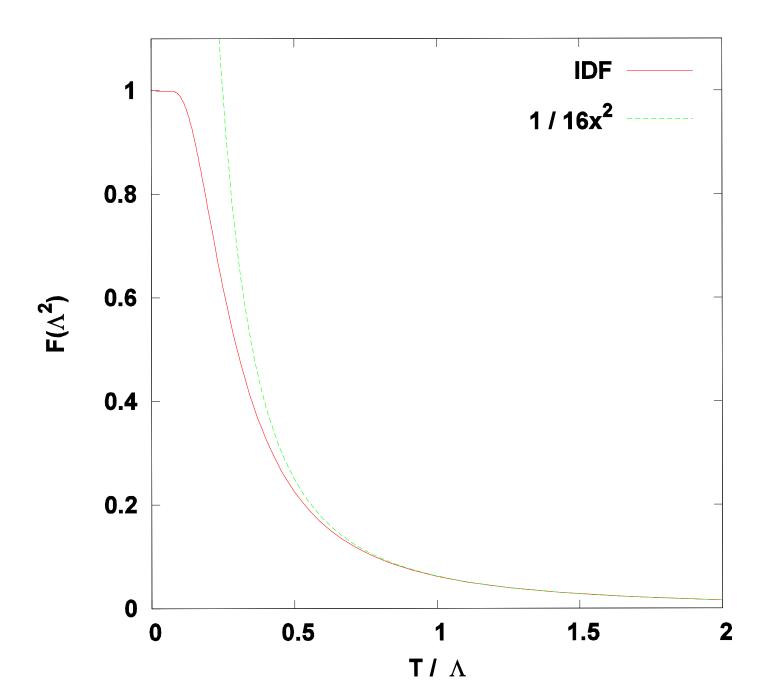
$$P(Q^{2}) = \frac{\iint dE_{1}dE_{2}d\theta E_{1}^{2}E_{2}^{2} e^{-\beta(E_{1}+E_{2})} \delta(Q^{2}-2E_{1}E_{2}(1-\cos\theta))}{\iint dE_{1}dE_{2}d\theta E_{1}^{2}E_{2}^{2} e^{-\beta(E_{1}+E_{2})}}$$





#### **Thermal expectation of NP order parameter**

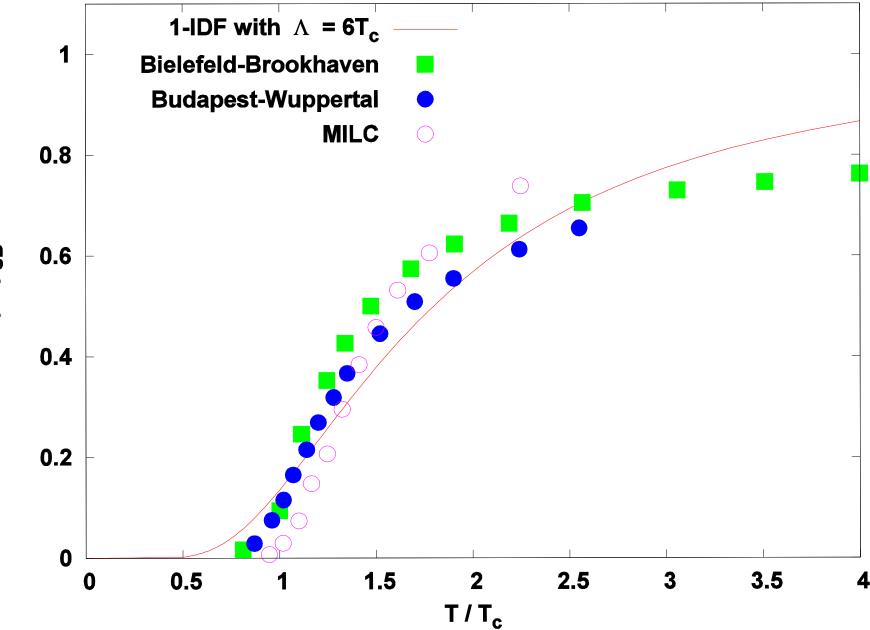




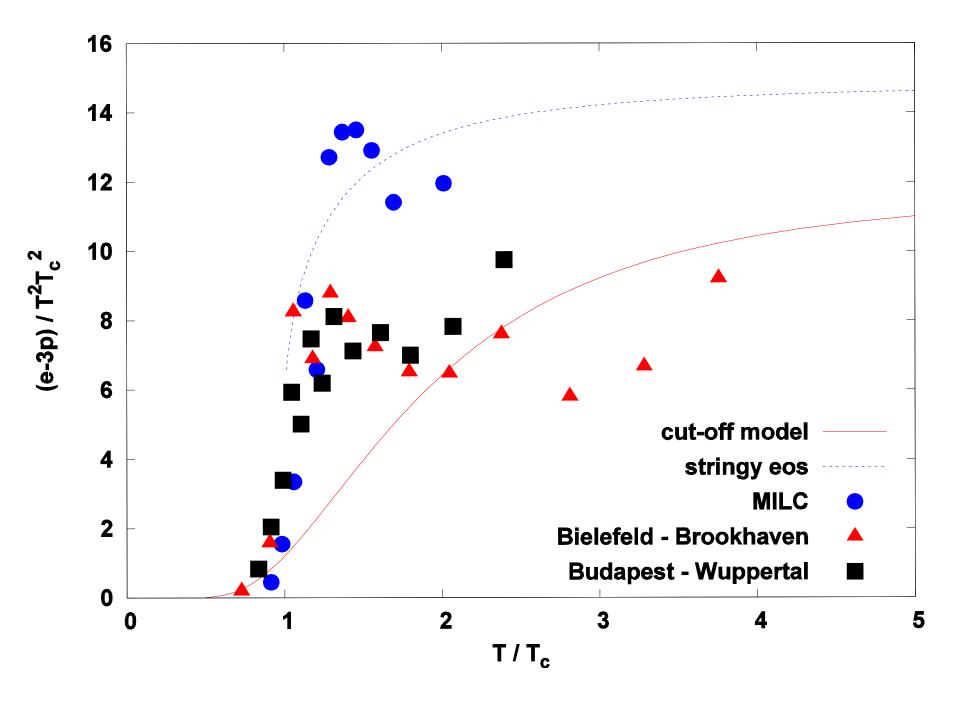
## NP effects at high T in the EoS

## $p = \frac{1}{3}\kappa T^4 - c_{NP}\Lambda^2 T^2$

 $e - 3p = 2c_{NP}\Lambda^2 T^2$ 



p / p<sub>SB</sub>



## To do list

- **1. Include massive quarks (strange)**
- Compare stringy matter and quasiparticle mass effects to leading order at high T
- 3. Compare stringy eos with power-law tail effects
- 4. Compare with other simple ideas, like Q2 cutoff

